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## STELLAR CHEMISTRY

# A New Way To Shine, A New Kind Of Star

by Staff Writers  
Cleveland OH (SPX) Dec 16, 2009

Dying, for stars, has just gotten more complicated. For some [stellar](#) objects, the final phase before or instead of collapsing into a black hole may be what a group of physicists is calling an electroweak star.



In their dying days, stars smaller than 2.1 times our Sun's mass die and collapse into neutron stars - objects dense enough that the neutrons and protons push against each other. More massive stars are thought to head toward collapse into a black hole.

Glenn Starkman, a professor of physics at Case Western Reserve University, together with former graduate students and post-docs De-Chang Dai and Dejan Stojkovic, now at the State University of New York in Buffalo, and Arthur Lue, at MIT's Lincoln Lab, offer a description of the structure of an electroweak star in a paper submitted to

Physical Review Letters Ordinary stars are powered by the fusion of light nuclei into heavier ones - such as hydrogen into helium in the center of our Sun.

Electroweak stars, they theorize, would be powered by the total conversion of quarks - the particles that make up the proton and neutron building blocks of those nuclei - into much lighter particles called leptons. These leptons include electrons, but especially elusive - and nearly massless - neutrinos.

"This is a process predicted by the well-tested Standard Model of [particle physics](#)," Starkman said. At ordinary temperatures it is so incredibly rare that it probably hasn't happened within the visible universe anytime in the last 10 billion years, except perhaps in the core of these electroweak stars and in the laboratories of some advanced alien civilizations, he said.

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But at the extreme temperatures and densities that might be reached when a star begins to collapse into a black hole, electroweak conversion of quarks into leptons should proceed at a rapid rate, the scientists say.

The energy generated could halt the collapse, much as the energy generated by [nuclear fusion](#) prevents ordinary stars like the Sun from collapsing.

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In other words, an electroweak star is the possible next step before total collapse into a black hole. If the electroweak burning is efficient, it could consume enough mass to prevent what's left from ever becoming a black hole.

Most of the [energy](#) eventually emitted from electroweak stars is in the form of neutrinos, which are hard to detect. A small fraction comes out as light and this is where the electroweak star's signature will likely be found, Starkman, said. But, "To understand that small fraction, we have to understand the star better than we do."

And until they do, it's hard to know how we can tell electroweak stars from other stars.

There's time, however, to learn. The theorists have calculated that this phase of a star's life can last more than 10 million years - a long time for us, though just an instant in the life of a star.

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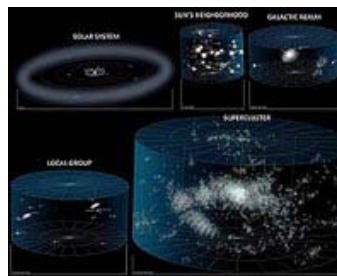
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## Astronomy Question Of The Week: Where Are The Nearest Islands Of Stars

Bonn, Germany (SPX) Dec 15, 2009

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